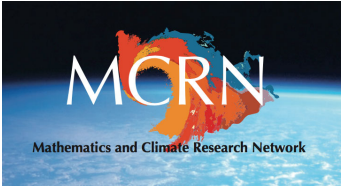
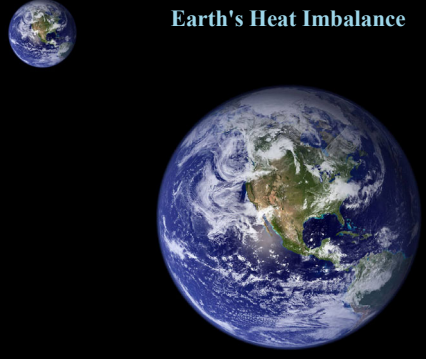


Earth's Heat Imbalance

Richard McGehee
School of Mathematics
University of Minnesota
Mathematics of Climate Seminar
September 10, 2020



Earth's Heat Imbalance



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Earth's Heat Imbalance

What Determines Earth's Surface Temperature?

"Energy Balance"

Conservation of Energy

temperature change \sim energy in $-$ energy out

short wave energy
from the Sun

long wave energy
from the planet

Everything else is detail.

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Earth's Heat Imbalance

Stefan-Boltzmann Law

$$F = \sigma T^4$$

power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

Reasonable approximation:
Every body in the solar system radiates energy according to this law.

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Earth's Heat Imbalance

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Example


surface temperature of the Sun: 5780K

power flux: $5.67 \times 10^{-8} \times (5780)^4 = 6.33 \times 10^7 \text{ W/m}^2$

total solar power output: $6.33 \times 10^7 \times 4\pi(r_s)^2$,
where $r_s =$ radius of the sun = $6.96 \times 10^8 \text{ m}$

total solar output: $3.85 \times 10^{26} \text{ W}$

230 nanoseconds = time it takes for the Sun to produce the equivalent of the annual global electricity production.



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Earth's Heat Imbalance

Insolation (Incoming solar Radiation)

Solar flux at a distance r from the sun:

$$F = \frac{6.33 \times 10^7 \times 4\pi r_s^2}{4\pi r^2} = 6.33 \times 10^7 \left(\frac{r_s}{r}\right)^2 \text{ W/m}^2$$

$r_s = 6.96 \times 10^8 \text{ m}$
 $r = 1.5 \times 10^{11} \text{ m}$


$F = 1368 \text{ W/m}^2$ ← solar flux at Earth's orbit

Power intercepted by the Earth: $F \times \pi r_e^2 \text{ W}$

Earth's surface area: $4\pi r_e^2 \text{ m}^2$

Average surface flux: $\frac{F \times \pi r_e^2}{4\pi r_e^2} = \frac{F}{4} = \boxed{342 \text{ W/m}^2}$

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Earth's Heat Imbalance

Insolation


Solar flux at a distance r from the sun:

$$F = \frac{6.33 \times 10^{27} 4\pi r_s^2}{4\pi r^2} = 6.33 \times 10^7 \left(\frac{r_s}{r}\right)^2 \text{ W/m}^2$$


$r_s = 6.96 \times 10^8 \text{ m}$
 $r = 1.5 \times 10^{11} \text{ m}$
 $F = 1368 \text{ W/m}^2$

Power intercepted by the Earth:
 $F \times \pi r_E^2 \text{ W}$, $r_E = \text{radius of Earth} = 6.37 \times 10^6 \text{ m}$
 $F = 1.74 \times 10^{17} \text{ W}$

Biologically Stored Energy
 total coal reserves: 10^{15} kg
 energy content: $3 \times 10^7 \text{ J/kg}$
 total energy in coal reserves: $3 \times 10^{22} \text{ J}$
= 2 days of insolation



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Earth's Heat Imbalance

Insolation

Global Average Insolation intercepted flux: $F = 1368 \text{ W/m}^2$
 Earth cross-section: πr_E^2
 surface area: $4\pi r_E^2$
 average flux: $1368/4 = 342 \text{ W/m}^2 = Q$

Simple Model

Assume that Earth is a perfectly thermally conducting black body.

$$Q = \sigma T^4$$


$$T = (Q/\sigma)^{1/4} = (342/5.67 \times 10^{-8})^{1/4}$$

$$= 279\text{K} = 6^\circ\text{C} = 43^\circ\text{F}$$

Dynamics
 $R \frac{dT}{dt} = Q - \sigma T^4$

heat capacity \rightarrow $R \frac{dT}{dt} = Q - \sigma T^4$ \rightarrow stable equilibrium

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Earth's Heat Imbalance

Albedo

Not all the insolation reaches the surface. Some is reflected back into space. The proportion reflected is called the albedo, denoted α .

For Earth, $\alpha \approx 0.3$.

Simple Model


Assume that Earth is a perfectly thermally conducting black body, but only 70% of the insolation is absorbed.

$$T = (0.7 \cdot F/\sigma)^{1/4} = (0.7 \cdot 342/5.67 \times 10^{-8})^{1/4}$$

$$= 255\text{K} = -18^\circ\text{C} = 0^\circ\text{F}$$

Dynamics
 $R \frac{dT}{dt} = Q(1 - \alpha) - \sigma T^4$ \rightarrow stable equilibrium

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Earth's Heat Imbalance

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Why isn't the Earth a Snowball?

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Earth's Heat Imbalance

The Greenhouse Effect

Greenhouse gases (CO_2 , H_2O , CH_4) are transparent to visible light, but opaque to infrared light. The energy from the sun passes through the atmosphere and heats the surface. The surface radiates energy at a lower temperature (infrared), which is absorbed by the atmosphere.

Who discovered the greenhouse effect?

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Earth's Heat Imbalance

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Who discovered the greenhouse effect?

Joseph Fourier (1827), Mémoire sur les Températures du Globe Terrestre et des Espaces Planétaires, *Mémoires de l'Académie Royale des Sciences*, t. vii., p. 569.




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Earth's Heat Imbalance


Why isn't the Earth a Snowball?

The Greenhouse Effect!

Joseph Fourier, *Mémoires de l'Académie des Sciences de l'Institut de France*, t. vii. 1827.



Svante Arrhenius, "On the Influence of Carbonic Acid in the Air upon the Temperature of the Ground," *Philosophical Magazine and Journal of Science (Fifth Series)* 41, pp. 237-276, 1896.

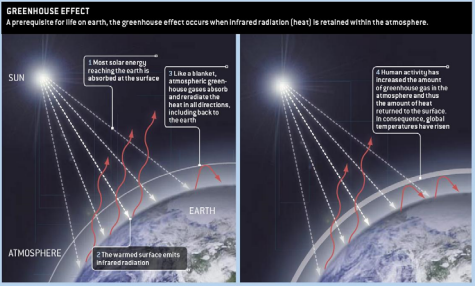


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Earth's Heat Imbalance

GREENHOUSE EFFECT

A prerequisite for life on earth, the greenhouse effect occurs when infrared radiation (heat) is retained within the atmosphere.

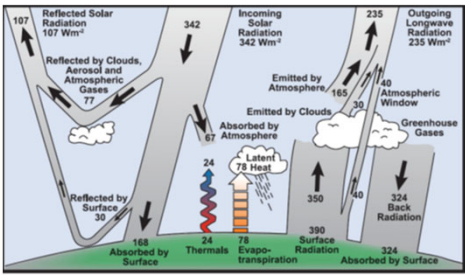


Gary Stix, *Scientific American* September 2006, pp.46-49

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Earth's Heat Imbalance

Heat Balance



Historical Overview of Climate Change Science, IPCC AR4, p.96
http://ipcc-wg1.ucar.edu/wg1/Report/AR4WG1_Print_CN01.pdf

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Earth's Heat Imbalance

OLR as a Function of Surface Temperature

(Outgoing Longwave Radiation)

$$OLR \approx A + BT$$

A and B are determined from satellite observations.
 T is surface temperature (in Celsius).

$$A = 202 \text{ W/m}^2$$

$$B = 1.90 \text{ W/m}^2\text{K}$$

Kelvin → Dynamics $R \frac{dT}{dt} = Q(1-\alpha) - \sigma T^4$ (photosphere temperature)

Celsius → becomes $R \frac{dT}{dt} = Q(1-\alpha) - (A + BT)$ (global mean surface temperature)

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Earth's Heat Imbalance

OLR as a Function of Surface Temperature

$$OLR \approx A + BT$$

Important:
 A+BT is not a linear approximation to the Stefan-Boltzmann equation.

Kelvin → Dynamics $R \frac{dT}{dt} = Q(1-\alpha) - \sigma T^4$ (photosphere temperature)

Celsius → becomes $R \frac{dT}{dt} = Q(1-\alpha) - (A + BT)$ (global mean surface temperature)

different

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Earth's Heat Imbalance

Homogeneous Earth

$$R \frac{dT}{dt} = Q(1-\alpha) - (A + BT)$$

Equilibrium Temperature: $Q(1-\alpha) - A - BT_{eq} = 0$

$$T_{eq} = \frac{Q(1-\alpha) - A}{B}$$

Stable, since $B > 0$.


Ice-free planet: $\alpha = 0.32$, $T_{eq} = 16^\circ\text{C}$
 Snowball planet: $\alpha = 0.62$, $T_{eq} = -38^\circ\text{C}$

No glacier would form on an ice-free Earth.
 No glacier would melt on a snowball Earth.

Easy question:
Why do we have ice caps?

Hard question:
If Earth was ever a snowball, how did we get out?

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Earth's Heat Imbalance

Latitude Dependence

Make T depend on $y = \sin(\text{latitude})$

$$R \frac{\partial T(y,t)}{\partial t} = Qs(y)(1-\alpha) - (A + BT(y,t))$$

insolation distribution

Q = global annual average insolation = 342 W/m^2

$s(y)$ = distribution across latitudes ($\int_0^1 s(y) dy = 1$)

One can show that


$$s(y) = \frac{2}{\pi^2} \int_0^{2\pi} \sqrt{1 - (\sqrt{1-y^2} \sin \beta \cos \theta - y \cos \beta)^2} d\theta$$

θ = obliquity = 23.4°

Chylek and Coakley's quadratic approximation:

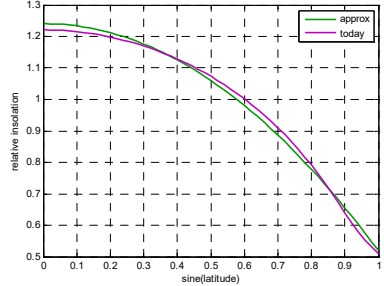
$$s(y) \approx 1 - 0.241(3y^2 - 1)$$

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Earth's Heat Imbalance


Insolation Distribution



green = quadratic approximation (Chylek & Coakley)

fuchsia = formula using obliquity of 23.4°

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Earth's Heat Imbalance

Latitude Dependence

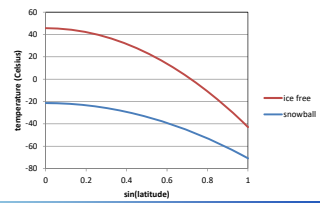
$$R \frac{\partial T(y,t)}{\partial t} = Qs(y)(1-\alpha) - (A + BT(y,t))$$

Note that y is just a parameter.


Equilibrium Temperature Profile

$$T_{eq}(y) = \frac{Qs(y)(1-\alpha) - A}{B}$$

$\alpha = 0.32$: ice free
 $\alpha = 0.62$: snowball

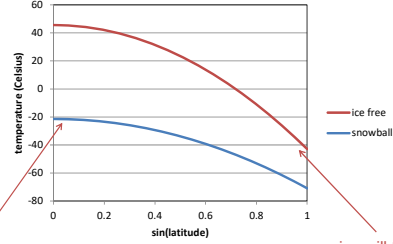


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Earth's Heat Imbalance


Latitude Dependence



ice won't melt (no exit from snowball)

ice will form (icecap)

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Earth's Heat Imbalance

What's Missing?

$$R \frac{\partial T(y,t)}{\partial t} = Qs(y)(1-\alpha) - (A + BT(y,t))$$

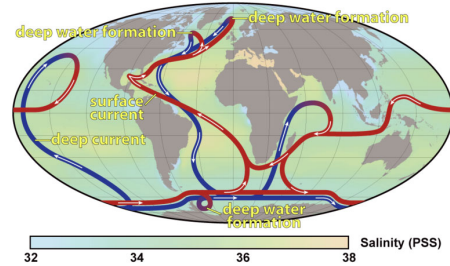
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Earth's Heat Imbalance

What's Missing?

Thermohaline Circulation



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Earth's Heat Imbalance

What's Missing?

Thermohaline Circulation

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Earth's Heat Imbalance

What's Missing?

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Earth's Heat Imbalance

What's Missing?

Thermohaline Circulation

Weather!

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Earth's Heat Imbalance

What's Missing?

<https://www.nytimes.com/2017/09/04/us/hurricane-irma.html>

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Earth's Heat Imbalance

What's Missing?

Weather!

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Earth's Heat Imbalance

Budyko's Model

$$R \frac{\partial T}{\partial t} = Qs(y)(1-\alpha) - (A + BT) + C(T - T_0)$$

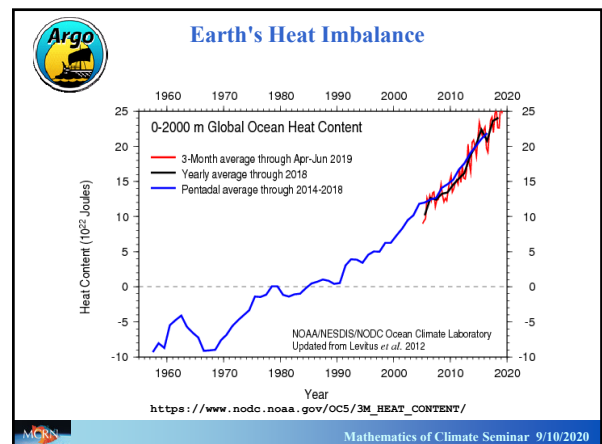
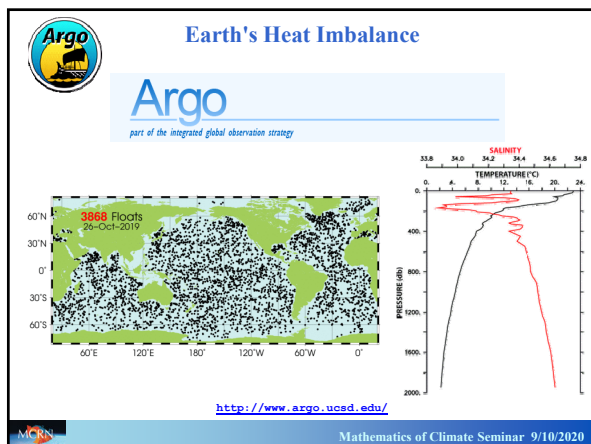
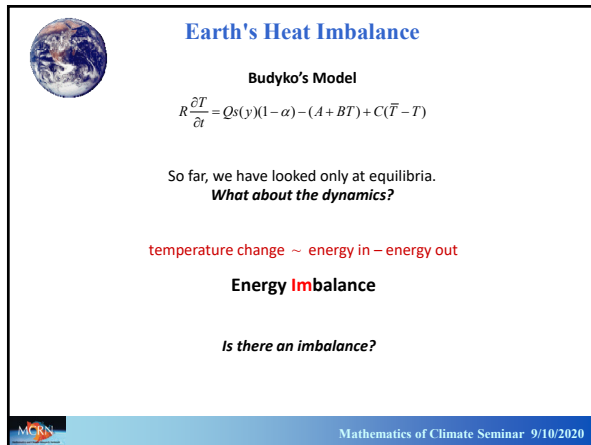
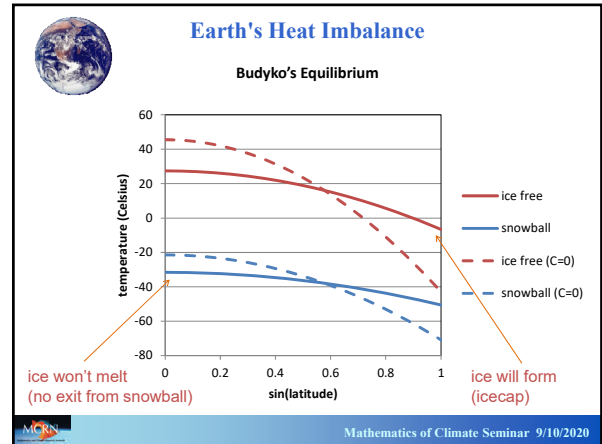
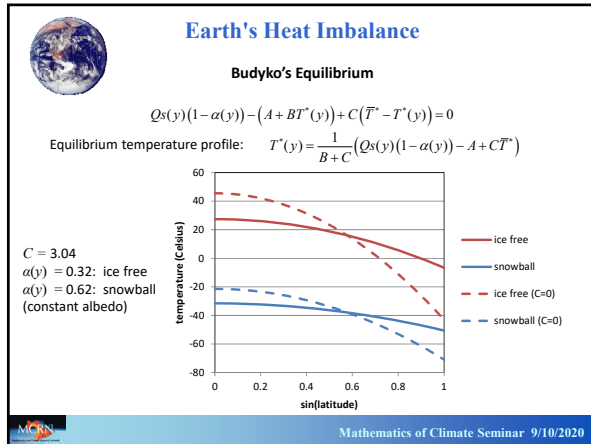
global mean temperature $\bar{T}(t) = \int_0^\pi T(y, t) dy$ **Weather**

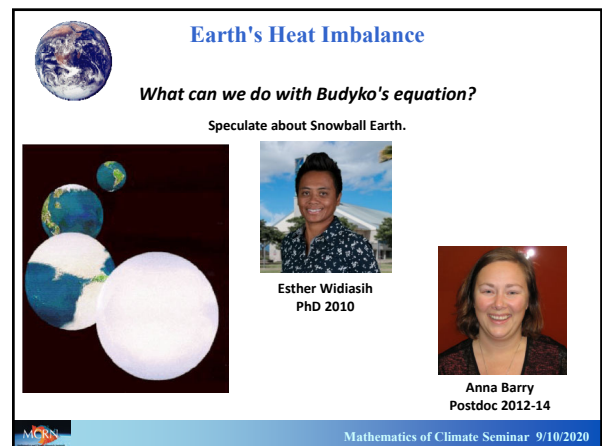
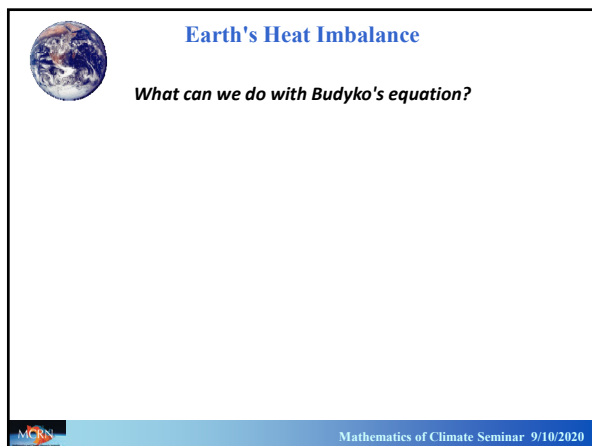
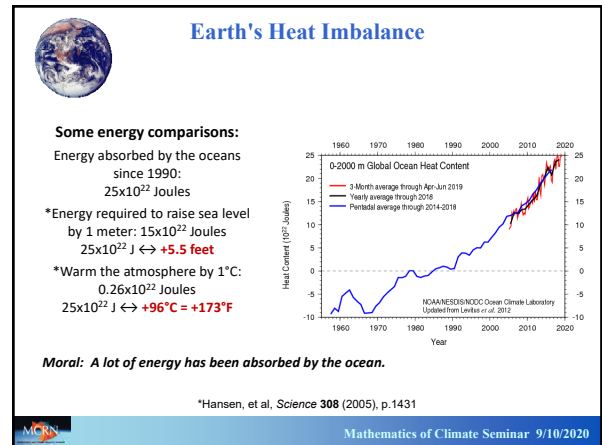
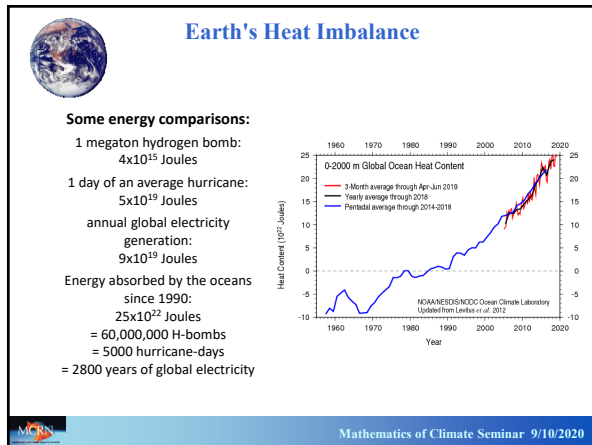
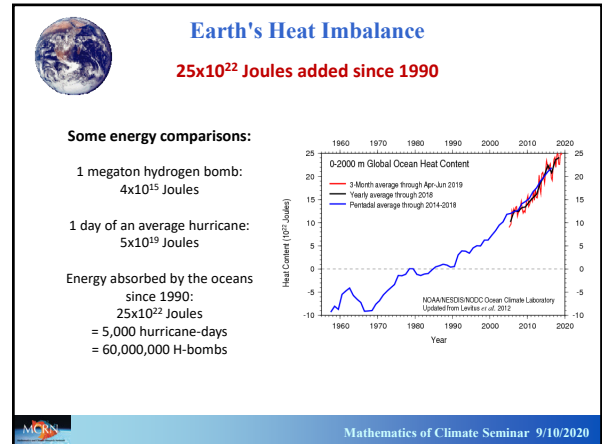
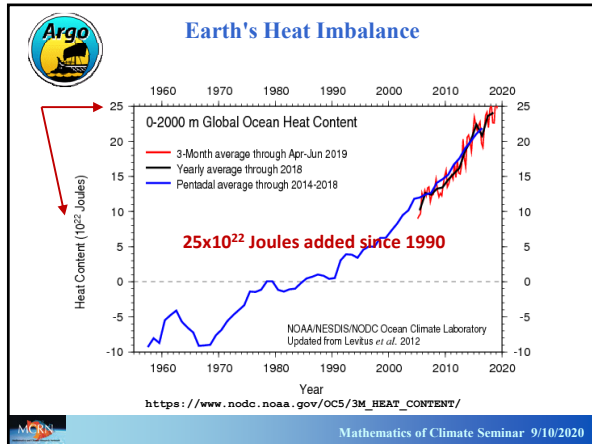
Second Law of Thermodynamics:
Energy travels from hot places to cold places.

Mikhail Budyko

Budyko's equation as a dynamical system:
 T lives in a function space (temperature as a function of latitude).

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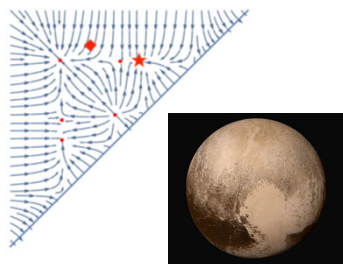





Earth's Heat Imbalance

What can we do with Budyko's equation?

Explain Pluto's heart.

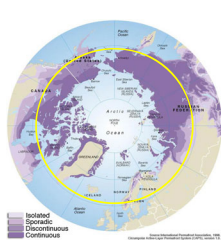
Alice Nadeau, PhD 2019

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Earth's Heat Imbalance

What can we do with Budyko's equation?

Predict permafrost melt.



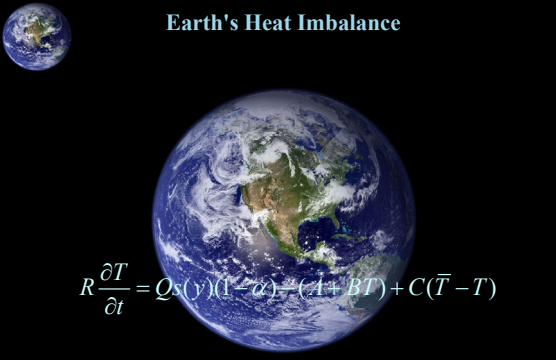
Aileen Zebrowski, BS 2016



Kaitlin Hill, Postdoc 2017-19

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Earth's Heat Imbalance



$$R \frac{\partial T}{\partial t} = Q_s(\gamma)(1 - \alpha) - (A + BT) + C(\bar{T} - T)$$

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